Mid IR Coherent Remote Sensing of Chemicals ... using QCLs, of course

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Science & Technology Facilities Council Rutherford Appleton Laboratory

Outline

- **Ø** Introduction
- Coherent (heterodyne) advantages
- **Ø** Historical perspective
- Coherent systems
 - Passive laser heterodyne radiometer
 - Active coherent laser spectrometer
 - Chirped laser dispersion spectrometer
- Conclusion





Introduction

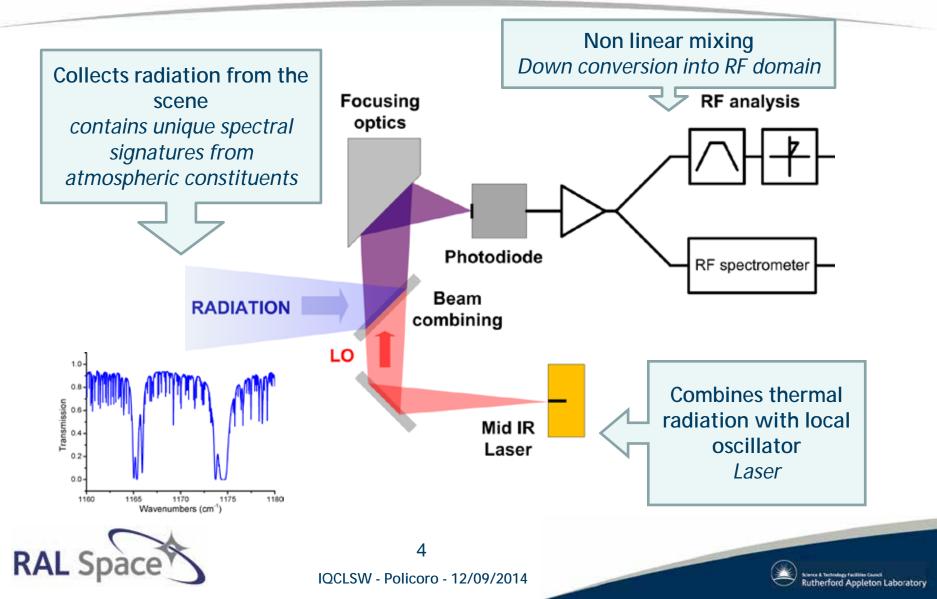
Observational requirements in remote sensing

- Remote sensing of our environment
 - Atmospheric composition information
 - Setting model constraints
 - Model validation
 - Ground-based, airborne, spaceborne
- Requirements for remote sounders
 - Driven by the need of fine meshes for models
 - Higher geographical resolution
 - Higher altitudinal resolution
 - Higher sensitivity and traceability





Optical Heterodyne Spectroscopy Basic principles



Advantages in Remote Sounding

Merits	Figures	Remote sounding benefits
High sensitivity Shot noise limited	<u>NEP = 4.10⁻¹⁶ W</u> (10µm -1s) <u>NESR = 120 nW/cm⁻².sr.cm⁻¹</u>	Detection of ultra-low concentration traces High accuracy - Heterodyne gain
High spectral resolution Set by electronic filters	<u>Resolving power > 10⁶</u> Resolution down to ~10 MHz Highest in the thermal IR	Full lineshape resolution Deconvolution of altitudinal information Usage of spectral micro-windows
High spatial resolution Coherent FoV	10 cm aperture gives FoV = 0.13 mrad = 27 arcsec Þ ~50 m LEO , ~4km GEO	Ultrafine geographical coverage Higher altitude resolution (limb) Less cloud interferences - Localized emission
Electrical definition of Instrument Lineshape	Directly measureable to a high level of accuracy	No ILS artefact ILS stability with sounding configuration
Miniaturization Enabled by QCLs	0.01 m3 – few kg – 30 W	Deployment on small satellites, piggy-backing, UAVs, HALE, ground networks
Phase information		Interferometry Dispersion measurements





Optical Laser Heterodyne Spectroscopy Historical Perspective (Mid IR)

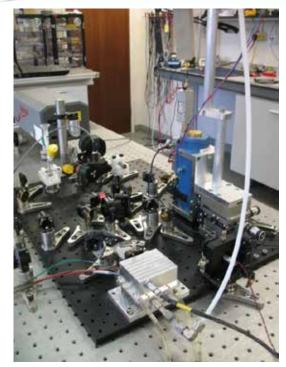
Ø High accuracy molecular resonance frequencies

- Atmospheric applications
 - Ground based observation (Menzies, 1974)
 - Pollution, stratospheric chemistry, ...
 - Space based planned (shuttle)
- Ground-based Astronomy
 - Titan stratospheric winds (Kostiuk, 2005)
 - Gain amplification on Mars and Venus (Mumma, 1981)
 - Shoemaker Levi collision with Jupiter
 - NH3 temporal behaviour (Fast 2005)
 - Mount Wilson infrared interferometer





Laser Heterodyne Spectro-Radiometer



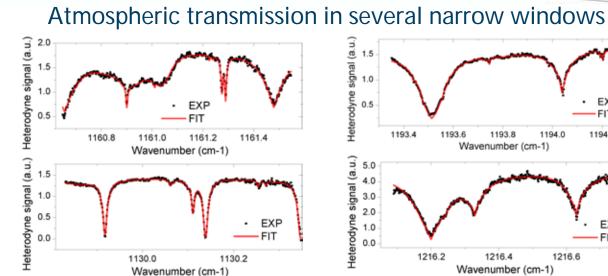


60 MHz (0.002 cm⁻¹)

1/40 solar disk

8-9 µm





Full lineshape information recovered

- contains the integrated altitudinal lineshapes

Narrow spectral windows (< 1cm⁻¹)

- can be optimized to increase information content
- limits interference
- Better control on error propagation

Faster retrievals

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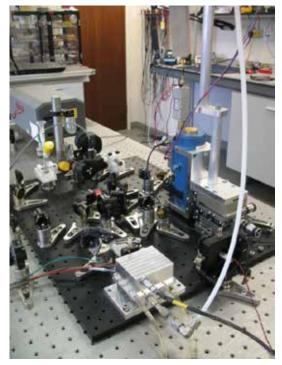


EXP

FIT

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Laser Heterodyne Spectro-Radiometer



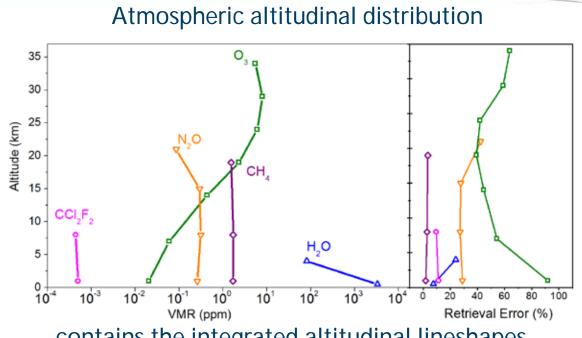


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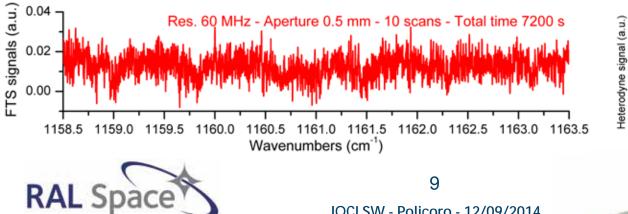
FTIR / LHR Side by Side Comparison Identical resolution 60 MHz and field of view

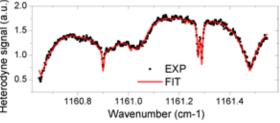


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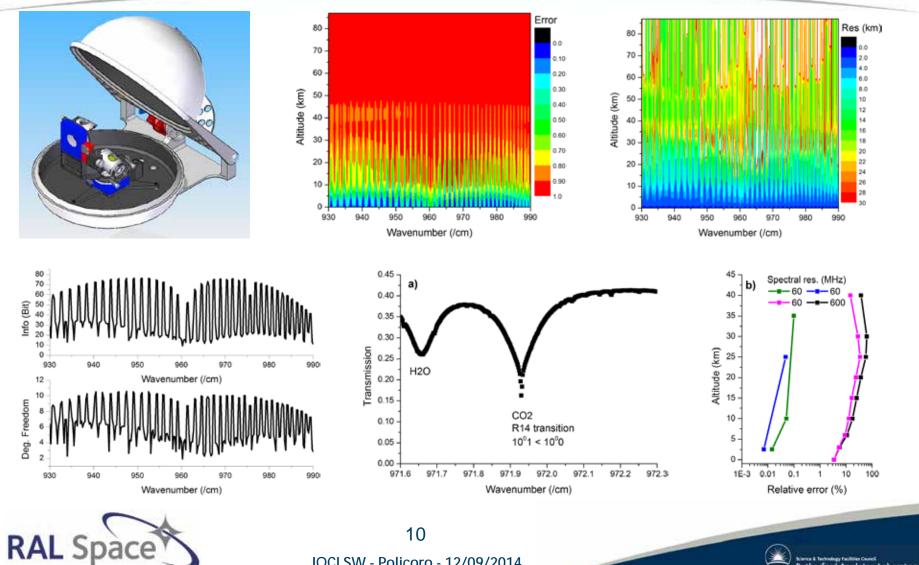
Bruker IFS 125HR - 4m x 2m

Bench top LHR - 1m² – 1min acquisition





Carbon Dioxide Sensing Programme On-going work on ground based LHR



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Miniaturization / Ruggedization Rationale

Ø Easing deployment in any field through

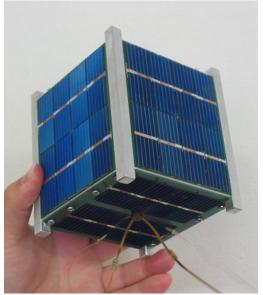
- Reduction of mass
- Reduction of volume
- Increased robustness
- **Ø** Ultimately cost reduction is the prime target
 - Launch between \$4k-\$30k per kilo
 - Up to few \$M for piggy-backing
 - Cost of space qualification

Also new environmental sounding opportunities

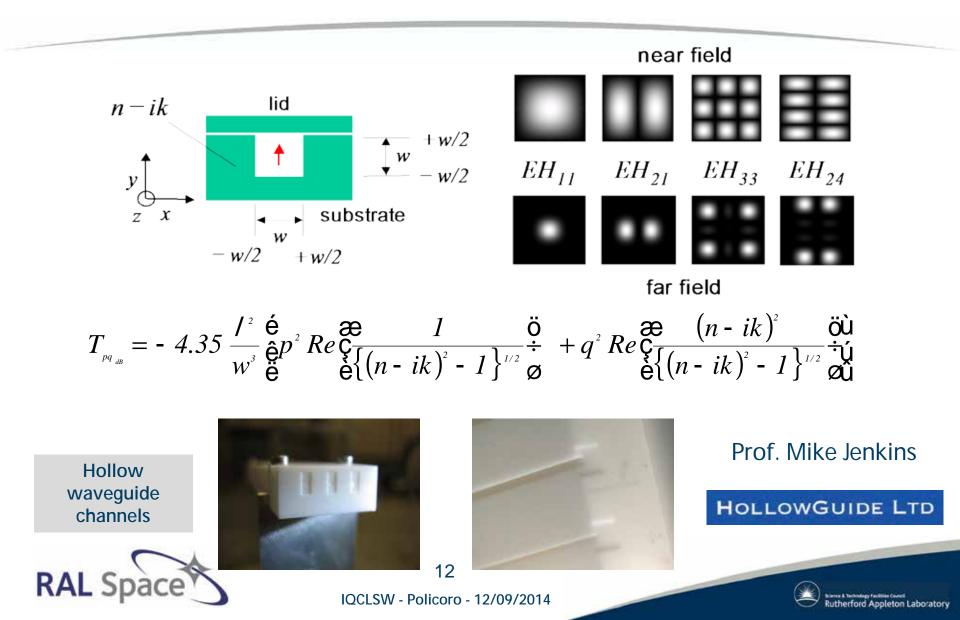
- Novel airborne platforms (UAVs, HALE)
- Development of micro- and nano-satellites (1-10 and 10-100 kg resp.)
- Piggy backing
- Dense ground network



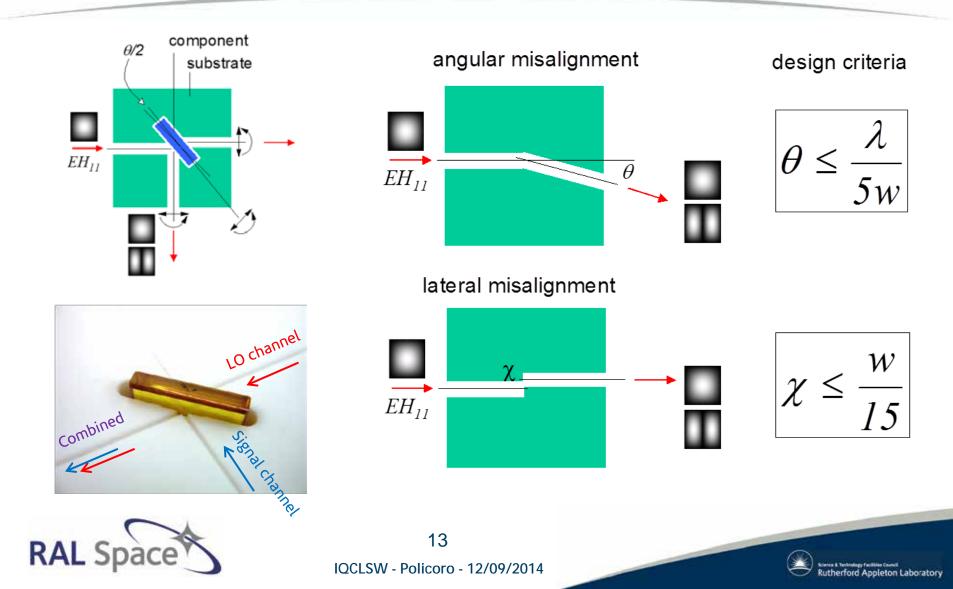




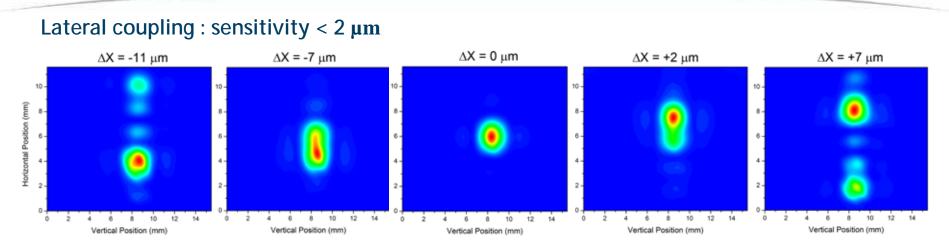
Multimode Hollow Waveguide



Optical Integration & Tolerancing

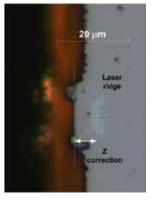


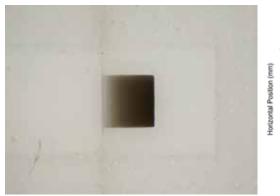
Optimum Coupling Assessment Far field profiles (0.750 mm guide width)



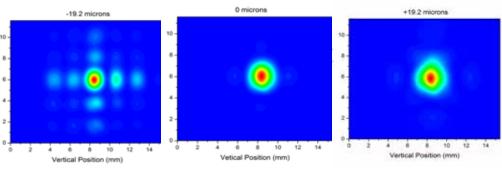
QCL output 4x10 µm²

HW input ~0.75x0.75 mm²





Waist position sensitivity < 10 μm

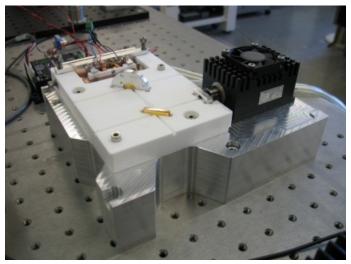


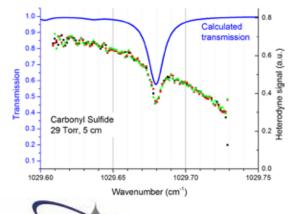




First Miniature LHR Demonstration Carbonyl sulphide absorption

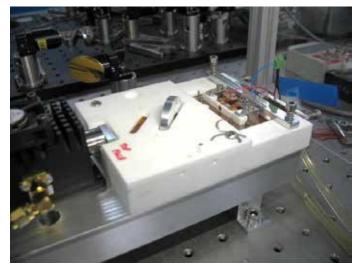
Iteration 1





RAL S

Iteration 2 – size further reduced



Level of performance achieved almost identical to "open space" traditional LHR using LN2 cooled detector

Still issue with the focusing on the detector -> understood and addressable



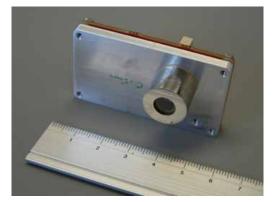
Miniaturization of Control Electronics QCL Driver – Photomixer

Miniaturization QCL control

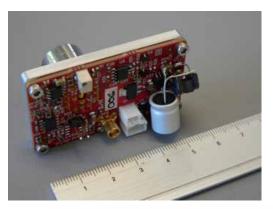
- 10 x 20 cm board
- While power efficiency X2
- While stability X4

Miniaturisation of detector electronics

T control and preamp











Active Version of LHR

Passive system well suited for long range TIR and FIR remote sounding

- Such as GEO, LEO, HAP
- Need thermal contrast
- For terrestrial activities such as
 - Highly localized emissions
 - Seeps of gas Emissions
 - Low concentration low vapour pressure
 - Short plume

An "active" version of the LHR was developed

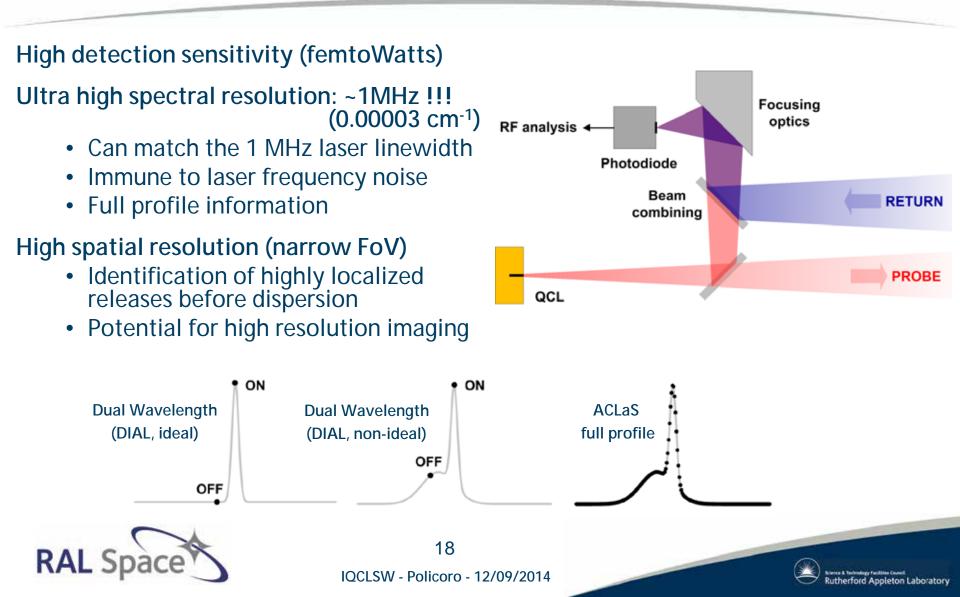
- Active Coherent Laser Spectrometer (ACLaS)
- Mid infrared spectroscopic heterodyne lidar



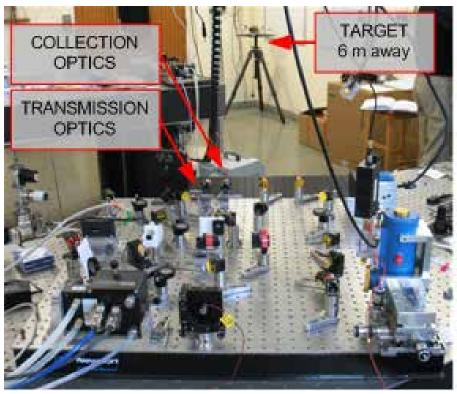


Benefits of ACLaS

Inherit advantages of LHR + new ones



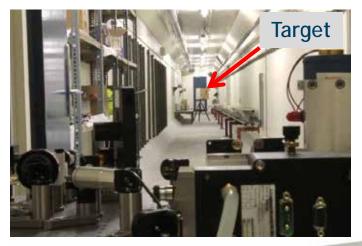
Bench Top Demonstrator of ACLaS



Short range tests 1 – 6 m

Medium range, up to 65 m

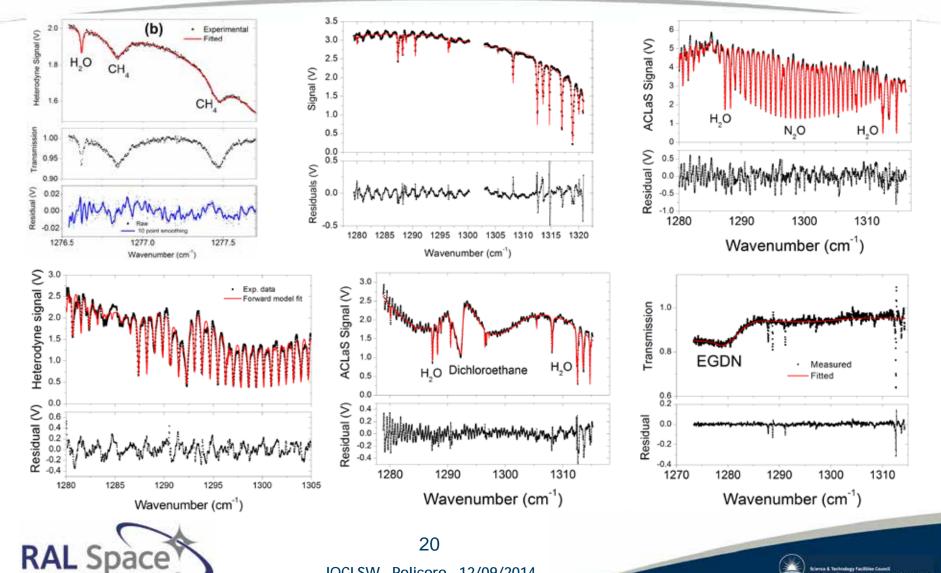








Some Results H2O, CH4, N2O, H2O2, DCE, EGDN



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ACLaS Limits of Detection Normalized to 1 m plume and 1s acquisition

Experimental data

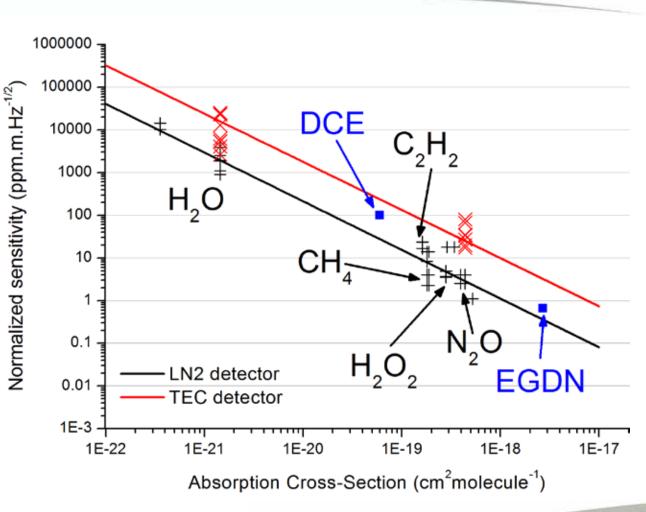
- 6 to 12 m standoff range
- ~20mW laser illumination
- EC-QCL degrades sensitivity by ~one order of magnitude

Projected improvements

200 mW still well eye safe
> 10X smaller LoD

10⁴ improvement SNR

- Compared to ultimate limit
- Speckle reduction needed





ACLaS Limits of Detection Normalized to 1 m plume and 1s acquisition

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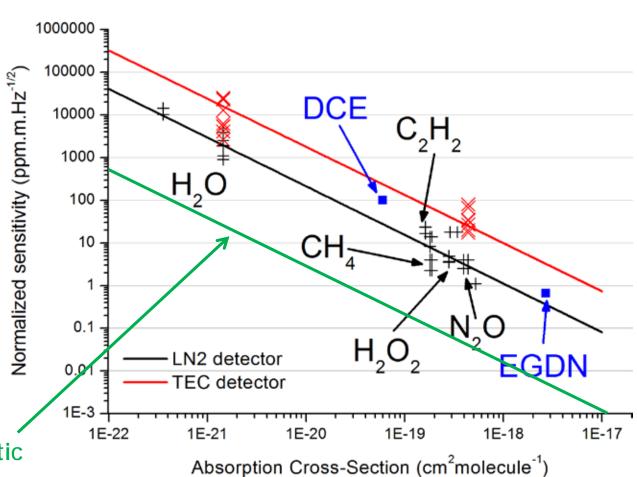
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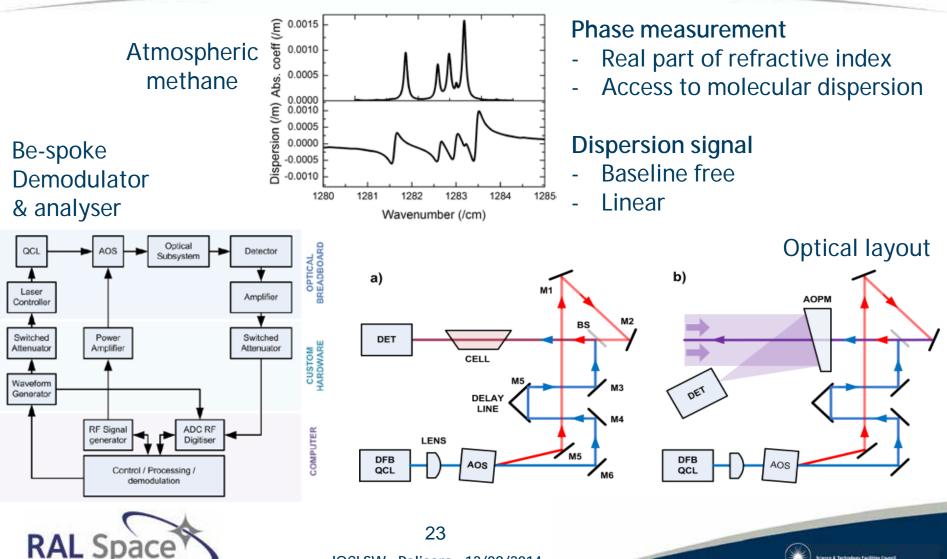
- Compared to ultimate limit
- Speckle reduction needed

100X improvement realistic



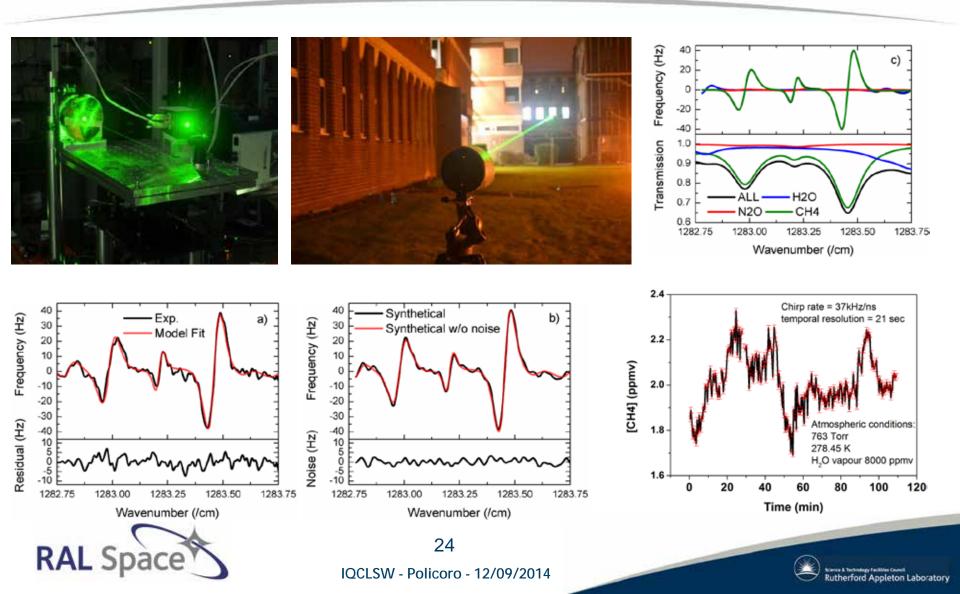


Measuring Molecular Dispersion Chirped Laser Dispersion Spectrometer (CLaDS)



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Open-Path Methane Monitoring



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